Konrad-Zuse-Zentrum für Informationstechnik Berlin



Adel Bouchareb, ¹ Chiang-Mei Chen, ² Gérard Clément, ³ Dmitri V. Gal'tsov, ⁴ Nikolai G. Scherbluk, ⁵ Thomas Wolf, ⁶

G_2 generating technique for minimal D = 5 supergravity and black rings

ZIB-Report 10-14 (July 2010)

¹ Laboratoire de Physique Théorique LAPTH (CNRS), Annecy-le-Vieux cedex, France

² Department of Physics, National Central University, Chungli 320, Taiwan

³ Laboratoire de Physique Théorique LAPTH (CNRS), Annecy-le-Vieux cedex, France

⁴ Department of Theoretical Physics, Moscow State University, 119899, Moscow, Russia

⁵ Department of Theoretical Physics, Moscow State University, 119899, Moscow, Russia

⁶ Department of Mathematics, Brock University, St.Catharines, Canada and ZIB Fellow

G_2 generating technique for minimal D = 5 supergravity and black rings

Adel Bouchareb^{*} and Gérard Clément[†] Laboratoire de Physique Théorique LAPTH (CNRS), B.P.110, F-74941 Annecy-le-Vieux cedex, France

Chiang-Mei Chen[‡] Department of Physics, National Central University, Chungli 320, Taiwan

Dmitri V. Gal'tsov[§] and Nikolai G. Scherbluk[¶] Department of Theoretical Physics, Moscow State University, 119899, Moscow, Russia

Thomas Wolf^{**}

Brock University, St. Catharines, Ontario, Canada L2S 3A1 (Dated: July 8, 2010)

A solution generating technique is developed for D = 5 minimal supergravity with two commuting Killing vectors based on the G_2 U-duality arising in the reduction of the theory to three dimensions. The target space of the corresponding 3-dimensional sigma-model is the coset $G_{2(2)}/(SL(2, R) \times SL(2, R))$. Its isometries constitute the set of solution generating symmetries. These include two electric and two magnetic Harrison transformations with the corresponding two pairs of gauge transformations, three SL(2, R) S-duality transformations, and the three gravitational scale, gauge and Ehlers transformations (altogether 14). We construct a representation of the coset in terms of 7×7 matrices realizing the automorphisms of split octonions. Generating a new solution amounts to transforming the coset matrices by one-parametric subgroups of $G_{2(2)}$ and subsequently solving the dualization equations. Using this formalism we derive a new charged black ring solution with two independent parameters of rotation.

PACS numbers: 04.20.Jb, 04.50.+h, 04.65.+e

I. INTRODUCTION

The discovery of rotating black rings [1] (for a recent review see [2] and references therein) has attracted new interest in five-dimensional minimal supergravity [3, 4]. Within this theory supersymmetric charged black ring solutions were found [5, 6]. The bosonic sector of five-dimensional minimal supergravity is Einstein-Maxwell theory with a Chern-Simons term, the structure of the Lagrangian being similar to that of eleven-dimensional supergravity [7, 8]. While in pure Einstein-Maxwell theory in five and higher dimensions no charged black hole solution, generalizing the uncharged Myers-Perry black holes [9], is known, the Chern-Simons term endows five-dimensional Einstein-Maxwell theory with more hidden symmetries, implying the existence of exact charged rotating black hole solutions [10–12], Meanwhile the most general black ring solution which might possess mass, two angular momenta, electric charge and magnetic moment as independent parameters is still not found. Here we propose a new generating technique which can solve this problem. It is based on the duality symmetries of the three-dimensional reduction of the theory.

The hidden symmetries arising upon dimensional reduction of five-dimensional minimal supergravity to three dimensions were studied by Mizoguchi and Ohta [7], by Cremmer, Julia, Lu and Pope [13] using the technique of [14], and were more recently investigated both in the bosonic and fermionic sectors by Possel [15] (see also [16]). The corresponding classical U-duality group is the non-compact version of the lowest rank exceptional group G_2 [17]. In three dimensions one obtains the gravity-coupled sigma-model with the homogeneous target space $G_{2(2)}/SO(4)$ for

^{*}Electronic address: bouchare@lapp.in2p3.fr

[†]Electronic address: gclement@lapp.in2p3.fr

[‡]Electronic address: cmchen@phy.ncu.edu.tw

[§]Electronic address: galtsov@phys.msu.ru

[¶]Electronic address: shcherbluck@mail.ru

^{**}Electronic address: twolf@BrockU.ca

the Lorentzian signature of the 3-space, or $G_{2(2)}/(SL(2, R) \times SL(2, R))$ in the Euclidean case. Some further aspects of these symmetries were discussed in [8], their infinite-dimensional extension upon reduction to two and one dimensions was also explored [18].

Here we investigate the $G_{2(2)}/(SL(2, R) \times SL(2, R))$ sigma model in the context of the solution generating technique which has proved to be extremely useful in various non-linear theories from pure gravity, Einstein-Maxwell theory [19–23] and dilatonic gravity [24–30] to more general supergravity models [31–33] and string theory [34]. Some partial use of hidden symmetries of this kind to generate new rotating rings recently became a rapidly developing industry. One direction was to use the SL(2, R) subgroup of the U-duality group [35, 36]. Another line was related to the purely gravitational sector (without the Maxwell field) which leads to SL(3, R) U-duality in three dimensions [20, 37, 38]. Further reduction to two dimensions gives rise to a Belinsky-Zakharov type integrable model which was extensively used to construct soliton solutions [38–49]. However, the full G_2 symmetry was never used for generating purposes for lack of a convenient representation of the coset $G_{2(2)}/(SL(2, R) \times SL(2, R))$ in terms of the target space variables. Although the 14-dimensional (adjoint) representation was given explicitly in [7], it is still too complicated for practical generating applications. Here we construct a suitable representation in terms of 7×7 matrices and give two examples of its application: a sigma-model construction of the electrically charged rotating black hole and the generation of a non-BPS doubly rotating charged black ring from the black ring with two angular momenta of [47].

Five-dimensional minimal supergravity contains a graviton, two N = 2 symplectic-Majorana gravitini (equivalent to a single Dirac gravitino), and one U(1) gauge field. The bosonic part of the Lagrangian is very similar to that of D = 11 supergravity, being endowed with a Chern-Simons term [3, 4]:

$$S_5 = \frac{1}{16\pi G_5} \left[\int d^5 x \sqrt{-\hat{g}} \left(\hat{R} - \frac{1}{4} \hat{F}^2 \right) - \frac{1}{3\sqrt{3}} \int \hat{F} \wedge \hat{F} \wedge \hat{A} \right],\tag{1}$$

where $\hat{F} = d\hat{A}$. This theory can be obtained as a suitably truncated Calabi-Yau compactification of D = 11 supergravity [50].

Our purpose is to construct a generating technique for classical solutions with two commuting Killing symmetries. Dimensional reduction leads to a three-dimensional sigma-model possessing a $G_{2(2)}$ target space symmetry [7]. To explore it fully we need a convenient representation of the action of symmetries on the target space variables. We give here an alternative derivation of the three-dimensional sigma-model which has the advantage of being more explicit and easy to use for solution generation. The reduction is performed in Sect. 2 in two steps, first to four, then to three dimensions. In Sect. 3 we reveal the symmetries of the three-dimensional sigma-model using a direct (computer assisted) solution of the corresponding Killing equations¹. The resulting symmetry transformations are identified in the usual terms of gauge, S-duality and Harrison-Ehlers sectors. Then we reformulate in Sect. 4 the problem in terms of a covariant (with respect to the two-Killing plane) reduction which is more suitable for constructing the matrix representation, and give the coset matrix representative as a symmetrical 7×7 matrix. In Sect. 5 we identify the charging transformation, and apply it to the construction of the doubly rotating charged black ring.

II. DIMENSIONAL REDUCTION

A. D=4

Assuming that the five-dimensional metric and the Maxwell field \hat{A} do not depend on a space-like coordinate z, we arrive at the four-dimensional Einstein theory with two Maxwell fields, a dilaton and an axion. We parametrize the five-dimensional interval and the Maxwell one-form as

$$ds_5^2 = e^{-2\phi} (dz + C_\mu dx^\mu)^2 + e^{\phi} ds_4^2, \qquad (2)$$

$$\hat{A} = A_{\mu}dx^{\mu} + \sqrt{3}\kappa dz, \tag{3}$$

 $(\mu = 1...4)$. The corresponding four-dimensional action reads

$$S_4 = \frac{1}{16\pi G_4} \int d^4x \sqrt{-g} \left[R - \frac{3}{2} (\partial\phi)^2 - \frac{3}{2} e^{2\phi} (\partial\kappa)^2 - \frac{1}{4} e^{-3\phi} G^2 - \frac{1}{4} e^{-\phi} \tilde{F}^2 - \frac{1}{2} \kappa F F^* \right],\tag{4}$$

¹ A purely algebraic construction of the Killing vectors will be presented elsewhere [51].

where

$$G_4 = G_5/2\pi R_5, \quad G = dC, \quad F = dA, \quad \tilde{F} = F + \sqrt{3}C \wedge d\kappa, \tag{5}$$

and F^* is the four-dimensional Hodge dual of F.

The dilaton ϕ and the axion κ parametrize the coset SL(2, R)/U(1). To reveal the SL(2, R) S-duality symmetry in the sector of vector fields (A is inherited from 5D theory, C is the Kaluza-Klein vector) one has to reparametrize them using some dualization [8]. We will reveal S-duality later on the level of the further 3D reduction.

The field equations in terms of the four-dimensional variables read

$$\nabla^2 \phi - e^{2\phi} (\partial \kappa)^2 + \frac{1}{4} e^{-3\phi} G^2 + \frac{1}{12} e^{-\phi} \tilde{F}^2 = 0, \qquad (6)$$

$$\nabla_{\mu} \left(e^{2\phi} \nabla^{\mu} \kappa \right) - \frac{1}{3} \left[\sqrt{3} \nabla_{\mu} (e^{-\phi} \tilde{F}^{\mu\nu} C_{\nu}) + \frac{1}{2} F_{\mu\nu} F^{*\mu\nu} \right] = 0, \tag{7}$$

$$\nabla_{\mu} \left(e^{-\phi} \tilde{F}^{\mu\nu} + 2\kappa F^{*\mu\nu} \right) = 0, \qquad (8)$$

$$\nabla_{\mu} \left(e^{-3\phi} G^{\mu\nu} \right) + \sqrt{3} e^{-\phi} \tilde{F}^{\mu\nu} \partial_{\mu} \kappa = 0, \qquad (9)$$

and the Bianchi identities are

$$\nabla_{\mu}F^{*\mu\nu} = 0, \qquad \nabla_{\mu}G^{*\mu\nu} = 0.$$
 (10)

It is convenient to introduce the modified Maxwell tensors

$$\mathcal{F}^{\mu\nu} = \mathrm{e}^{-\phi} \tilde{F}^{\mu\nu} + 2\kappa F^{*\mu\nu},\tag{11}$$

$$\mathcal{G}^{\mu\nu} = \mathrm{e}^{-3\phi} G^{\mu\nu} + \sqrt{3} \left(\mathrm{e}^{-\phi} \kappa \tilde{F}^{\mu\nu} + \kappa^2 F^{*\mu\nu} \right), \tag{12}$$

in terms of which the Maxwell equations have the divergence form

$$\nabla_{\mu}\mathcal{F}^{\mu\nu} = 0, \qquad \nabla_{\mu}\mathcal{G}^{\mu\nu} = 0. \tag{13}$$

B. D=3

Further reduction to D = 3 is performed with respect to time, assuming the standard parametrization of the stationary four-metric

$$ds_4^2 = -f(dt - \omega_i dx^i)^2 + f^{-1}h_{ij}dx^i dx^j.$$
(14)

The spatial part of the Bianchi identities (10) is solved introducing the electric potentials $C_0 = \bar{v}_1$, $A_0 = \bar{v}_2$, so that

$$G_{i0} = \partial_i \bar{v}_1, \qquad F_{i0} = \partial_i \bar{v}_2. \tag{15}$$

Similarly, the spatial components of the Maxwell equations (13) are solved by introducing magnetic potentials \bar{u}_1 , \bar{u}_2 :

$$\mathcal{G}^{ij} = \frac{f}{\sqrt{h}} \epsilon^{ijk} \partial_k \bar{u}_1, \qquad \mathcal{F}^{ij} = \frac{f}{\sqrt{h}} \epsilon^{ijk} \partial_k \bar{u}_2. \tag{16}$$

The time components of the corresponding equations then give the second order equations for these potentials. Straightforwardly we can find (with the convention $\epsilon^{ijk} = -\epsilon^{0ijk}$)

$$G^{ij} = \frac{f}{\sqrt{h}} e^{3\phi} \epsilon^{ijk} (w_1)_k, \qquad (w_1)_k := \partial_k \bar{u}_1 - \sqrt{3} \kappa (\partial_k \bar{u}_2 - \kappa \partial_k \bar{v}_2), \tag{17}$$

$$\tilde{F}^{ij} = \frac{f}{\sqrt{h}} e^{\phi} \epsilon^{ijk} (w_2)_k, \qquad (w_2)_k \coloneqq \partial_k \bar{u}_2 - 2\kappa \partial_k \bar{v}_2, \tag{18}$$

$$\tilde{F}_{i0} = (z_2)_i, \qquad (z_2)_i := \partial_i \bar{v}_2 - \sqrt{3} \bar{v}_1 \partial_i \kappa. \qquad (19)$$

The remaining components of the Maxwell tensors are obtained using the following relations valid for any second rank four-dimensional antisymmetric tensor $W_{\mu\nu}$ with the assumed form of the metric (14):

$$W^{i0} = W^{ij}\omega_j - h^{ij}W_{j0}, \qquad W_{ij} = f^{-2}h_{ik}h_{jl}W^{kl} + 2W_{0[i}\omega_{j]}.$$
(20)

Using this we find:

$$G^{i0} = \frac{f}{\sqrt{h}} e^{3\phi} \epsilon^{ijk} \omega_j(w_1)_k - \partial^i \bar{v}_1, \qquad (21)$$

$$G_{ij} = f^{-1}\sqrt{h}e^{3\phi} \epsilon_{ijk}(w_1)^k + 2\omega_{[i}\partial_{j]}\bar{v}_1, \qquad (22)$$

$$\tilde{F}^{i0} = \frac{f}{\sqrt{h}} e^{\phi} \epsilon^{ijk} \omega_j(w_2)_k - (z_2)^i, \qquad (23)$$

$$\tilde{F}_{ij} = f^{-1} \sqrt{h} e^{\phi} \epsilon_{ijk} (w_2)^k + 2\omega_{[i} (z_2)_{j]}, \qquad (24)$$

and for the squared quantities:

$$G^{2} = -2(\partial \bar{v}_{1})^{2} + 2e^{6\phi}(w_{1})^{2}, \qquad \tilde{F}^{2} = -2(z_{2})^{2} + 2e^{2\phi}(w_{2})^{2}, \qquad (25)$$

where $(w_1)^2 = (w_1)_i (w_1)^i$.

Now we turn to the Einstein equations:

$$R_{\mu\nu} - \frac{3}{2} \left(\partial_{\mu}\phi \partial_{\nu}\phi + e^{2\phi} \partial_{\mu}\kappa \partial_{\nu}\kappa \right) - \frac{1}{2} e^{-3\phi} \left(G_{\mu\alpha} G_{\nu}{}^{\alpha} - \frac{1}{4} G^2 g_{\mu\nu} \right) - \frac{1}{2} e^{-\phi} \left(\tilde{F}_{\mu\alpha} \tilde{F}_{\nu}{}^{\alpha} - \frac{1}{4} \tilde{F}^2 g_{\mu\nu} \right) = 0.$$
(26)

The Ricci tensor decomposes as follows

$$R_{00} = \frac{1}{2} \left(f \nabla^2 f - (\partial f)^2 + \tau^2 \right), \qquad (27)$$

$$R^{i}{}_{0} = -\frac{f}{2\sqrt{h}} \epsilon^{ijk} \partial_{j} \tau_{k}, \qquad (28)$$

$$R^{ij} = f^2 \mathcal{R}^{ij} - \frac{1}{2} \left[\partial^i f \partial^j f + \tau^i \tau^j \right] + h^{ij} R_{00}, \qquad (29)$$

where

$$\tau^{i} = -\frac{f^{2}}{\sqrt{h}} \epsilon^{ijk} \partial_{j} \omega_{k}.$$
(30)

The 0*i*-part of (26) can be solved introducing the twist potential χ via

$$\tau_i = \partial_i \chi + \frac{1}{2} \left\{ \bar{v}_1 \partial_i \bar{u}_1 - \bar{u}_1 \partial_i \bar{v}_1 + \bar{v}_2 \partial_i \bar{u}_2 - \bar{u}_2 \partial_i \bar{v}_2 + \sqrt{3} \left[\kappa^2 \bar{v}_1 \partial_i \bar{v}_2 - \bar{v}_2 \partial_i (\kappa^2 \bar{v}_1) \right] - \sqrt{3} \left[\kappa \bar{v}_1 \partial_i \bar{u}_2 - \bar{u}_2 \partial_i (\kappa \bar{v}_1) \right] \right\}.$$
(31)

Using this relation, we can rewrite the 00-component of the Einstein equations as

$$R_{00} = \frac{1}{4} f \left\{ e^{-3\phi} \left[(\partial \bar{v}_1)^2 + e^{6\phi} (w_1)^2 \right] + e^{-\phi} \left[(z_2)^2 + e^{2\phi} (w_2)^2 \right] \right\},$$
(32)

and present the space-space part as

$$R^{ij} = \frac{3}{2} f^2 h^{ia} h^{jb} \left(\partial_a \phi \partial_b \phi + e^{2\phi} \partial_a \kappa \partial_b \kappa \right) - \frac{1}{2} e^{-3\phi} f \left[\partial^i \bar{v}_1 \partial^j \bar{v}_1 + e^{6\phi} (w_1)^i (w_1)^j \right] + \frac{1}{4} e^{-3\phi} f h^{ij} \left[(\partial \bar{v}_1)^2 + e^{6\phi} (w_1)^2 \right] - \frac{1}{2} e^{-\phi} f \left[(z_2)^i (z_2)^j + e^{2\phi} (w_2)^i (w_2)^j \right] + \frac{1}{4} e^{-\phi} f h_{ij} \left[(z_2)^2 + e^{2\phi} (w_2)^2 \right].$$
(33)

From the Eq. (28) we then find that the Ricci tensor built on the three-dimensional metric h_{ij} will satisfy the following three-dimensional Einstein equation

$$\mathcal{R}_{ij} = \frac{1}{2f^2} \left(\partial_i f \partial_j f + \tau_i \tau_j \right) + \frac{3}{2} \left(\partial_i \phi \partial_j \phi + e^{2\phi} \partial_i \kappa \partial_j \kappa \right) - \frac{1}{2f} \left[e^{-3\phi} \partial_i \bar{v}_1 \partial_j \bar{v}_1 + e^{3\phi} (w_1)_i (w_1)_j + e^{-\phi} (z_2)_i (z_2)_j + e^{\phi} (w_2)_i (w_2)_j \right].$$
(34)